Laboratory Tests of Gravity Workshop Eöt-Wash Laboratory Jan 5. 2007

Schedule:

- 1:00 Welcome and introductions
- 1:10 Overview of lab tests of gravity (Eric Adelberger)
- 1:25 Tests of the gravitational inverse-square law (Eric Adelberger)
- 1:45 Tests of Einstein's Equivalence Principle (Stephan Schlamminger)
- 2:05 Measurements of Newton's constant G (Jens Gundlach)
- 2:25 First laboratory tour
- 2:55 Coffee, cookies and discussion
- 3:20 Lunar Laser Ranging (James Battat)
- 3:40 LISA gravitational wave detector (Scott Pollack)
- 4:00 Preferred frame (Lorentz invariance) tests (Claire Cramer)
- 4:20 Second lab tour
- 5:00 Bus arrives to pick up participants



 Problems in unifying gravity with the other forces in physics have convinced many of us that we must be missing something BIG.

- Important to test cherished concepts for possible clues about what that could be
- 1/r² law
- Equivalence Principle
- Preferred frames?



WMAP view of the Anisotropy of the Cosmic Microwave Background





WMAP anisotropy before removing the effects of our velocity through the CMB





There really is a preferred frame set by the universe as a whole! AAPT Workshop, Seattle January 2007

The Eöt-Wash[®] group in experimental gravitation

Faculty EGA Jens Gundlach Blayne Heckel Staff Erik Swanson Postdocs Seth Hoedl CD Hoyle Stephan Schlamminger Current Grad students Claire Cramer Ted Cook Charlie Hagedorn William Terrano Todd Wagner





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WHY IS GRAVITY SO WEAK? Comparison of electrical & gravitational attractions of e and p at a given separation r. $F_e = \frac{1}{4\pi\epsilon_b} \frac{9_1 9^a}{T^2} \qquad F_g = 6 \frac{m_1 m_2}{r^2}$ $\frac{F_e}{F_g} = \frac{2 \times 10^{39}}{r_m m^2}$ WHY IS THE COSMOLOGICAL CONSTANT SO SMALL? • Einstein's biggest blunder? • Ciscovery of cosmic acceleration (Type 1A supernovae)

FAMOUS PROBLEMS WITH GRAVITY

• particle-physics predictions for Λ (vacuum energy) ~ 10¹²⁰ larger than observed value! ~ 10⁶⁰ if supersymmetry is just around the

corner

WHAT IS THE DARK MATTER?

we know that most of the "gravitational" effects in the Universe are not produced by normal matter





Any given test of the 1/12 law is sensitive to a restricted range of length scales



 $\frac{T_{A}^{2}}{r_{A}^{3}} = \frac{T_{B}^{2}}{r_{B}^{3}} ?$



precession of perigee?

... need many different approaches to cover a wide range of length scales



95% confidence limits as of 2000





Does dark energy define a new fundamental length scale in physics?

$$ho_{\rm d} \approx 3.8 \ {\rm keV/cm^3}$$

 $ho_{\rm d} = \sqrt[4]{\hbar c/\rho_{\rm d}} \approx 85 \ \mu{\rm m}$

a second "Planck length"?



95% confidence limits as of 2000





the Irvine experiment





AA







The hierarchy problem

 $\frac{mass \ scale \ of \ gravity}{V_{N}(r) = 6} \frac{m \ m}{r} = \frac{K_{C}}{M_{p}^{3}} \frac{m \ m}{r}$ $\rightarrow M_{p} = \sqrt{K_{C}/6} \sim 10^{16} \ TeV$

- mass scale of particle physics M_{SM} ~ 1 TeV
- Avkani-Hamed et al. solution to the problem
 Phys. Lett. B 429 (1998) 263

assume that:

- · gravity propagates in all of the 7 extra climensions of string theory
- · SM particles are confined to a 4-dim "brane"
- Some of the 7 extra dimensions are "large" while the remainder are "carled up" at the Planck scale $R_p = \sqrt{Gk/c^3} = 1.6 \times 10^{-33} cm$





Gauss's Law and extra dimensions





illustration from Savas Dimopoulis





Why might gravity get weak at separations less than 0.1 mm?

- the repulsive "gravity" deduced from cosmological data indicates that empty space has an energy of p = 4 kev/cm³.
- this corresponds to a length scale d= Vhclp ~ 0.1mm
- Sundrum's suggestion: the graviton string has a size of O.Imm. this prevents it from "seeing" the short-distance physics that produces most of the predicted Vacuum energy
- prediction: gravity gets very wak at separations less than 0.1 mm







The original Eöt-Wash instrument for testing the ISL at short length scales

Advantages:

planar geometry

can place rigid conductor between detector & attractor

signal $\omega \neq$ disturbance ω

approximate null for Newtonian gravity with no cancellation of new shortrange physics

twist readout insensitive to pendulum & wobble modes



the 10-hole pendulum



AAPT Workshop, Seattle January 2007

PhD project of CD Hoyle



the Stanford ISL Experiment using low-temperature micro-cantilevers



95% confidence-level constraints as of 2004





the 42-hole pendulum



tungsten fiber, 20µm diameter, 80cm length
leveling mechanism
3 aluminum calibration spheres
4 mirrors for tracking angle of deflection detector: 1mm thick molybdenum ring
with 42 holes arranged in 21-fold rotational symmetry
not pictured, 10µm thick Au-coated BeCu membrane, electrostatic shield

attractor : **rotating pair** of discs with 21fold rotational symmetry, holes in lower attractor out of phase with holes in upper attractor to **cancel Newtonian gravity**



PhD project of Dan Kapner AAPT Workshop, Seattle January 2007

Dan Kapner assembling the 42-hole instrument





It doesn't show, but he had a beard also



these data Were taken with calibration turn-table stationary







raw signal

2-pt digital filter used in our 10-hole work

5-pt digital filer

combined analysis of all 3 experiments





some "gee-whiz" numbers

- typical error corresponds to light spot on detector moving by 0.6 nm
- typical torque in our 42-hole experiments is ~ 1fN-m with statistical uncertainty of ~0.006 fN-m
- corresponds to a force ~(40±0.24) fN
- suppose you could cut a postage stamp into 10¹² equal pieces
- typical force is 60 times the weight of 1 of those pieces
- typical statistical error is ~1/3 the weight of 1 piece



the Fourier-Bessel pendulum







will be the PhD project of Ted Cook







Velocity = Mass x Acceleration

Mathey-Tissot"

The time has come to greet exhilaration and accomplishment at the bottom of this mountain. Decades of experience have lead you to the edge. Each moment must be precise and confident. At this point there is one direction; forward.

